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S/048/60/024/009/009/015 B0:3/B063

24.6720

AUTHORS:

Surkov, Yu. A., Chernov, G. M., Lavrukhina, A. K.,

KIN OMONOMKOPEB. V.

TITLE:

Investigation of Meutron-deficient Osmium Isotopes  $oldsymbol{eta}$ 

PERIODICAL

Izvestiya Akademii nauk SSSR. Seriya fizicheskaya, 1960,

Vol. 24, No. 9, pp. 1119-1123

TEXT: The present paper gives the results of an investigation of neutron-deficient osmium isotopes on the synchrocyclotron of OIYsI (Joint Institute of Nuclear Research). The osmium isotopes were produced by bombarding  $\sim 0.2$  g of gold with 660-Mev protons for 1  $\sim 2$  hours. The purity of the separated elements was radiochemically checked. A 100-channel scintillation gamma spectrometer and  $\beta$ - and  $\gamma$ -counters were used to analyze the beta and gamma emission of the nuclear reaction products. Fig. 1 shows the gamma spectrum of the osmium isotopes obtained. The authors identified spectrum of the osmium isotopes obtained. The authors identified of at least 10 hours. Os 182, Os 183 and Os which had a half-life of at least 10 hours. In addition, the spectrum showed an intense line, Ey = 230 keV, which had

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Investigation of Neutron-deficiers Osmium Isotopes

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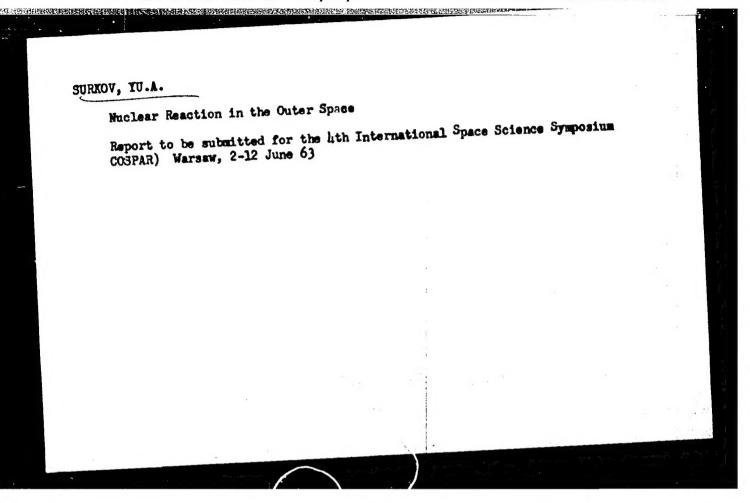
a half-life of ~2.7 hours. Control experiments indicated the existence of a new osmium isotope having a half-life of three hours. This was confirmed by the study of the daughter osmium (Figs. 3 and 4) and the daughter. rhenium (Fig. 5). This neutron-deficient isotope is assumed to be Os 181 and has a half-life 2.7 hours. By capture of the orbital electron it is converted into Re 181. 230-kev gamma quanta are emitted during this conversion. Fig. 2 shows the descending curve of the activity sum of Os, which was measured by an end-window counter. It confirms the correctness of the identification of the isotopes. As there are now only few data available on neutron-deficient Ir, Os, and Re isotopes, the conclusions drawn from the results obtained require further confirmation. The agreement of these results with experiments recently carried out with protons of 10 - 80 Mev (Ref. 1) indicates that the above-mentioned identification is correct. The results further indicate the existence of the isotope Ir 183 % which is formed by the decay of the two isomers  $0s^{183}$  and  $0s^{183}$ . Furthermore, a 23-min activity of osmium (E<sub>x</sub> = 170 keV) was observed during the experiments. However, the 23-min isotope may well be an isomer

ALIMARIN, I.P.; YAKOVLEY, Yn.V.; SHCHULEFNIKOV, M.N.; VIASOV, D.A.;
CHERNOV, G.M.; SUNKOV, Yn.A.

Radioactive determination of impurities in high purity
thallium. Zhur.amal.khim. 16 no.2:213-216 Mr-Ap '61.
(MIRA 14:5)

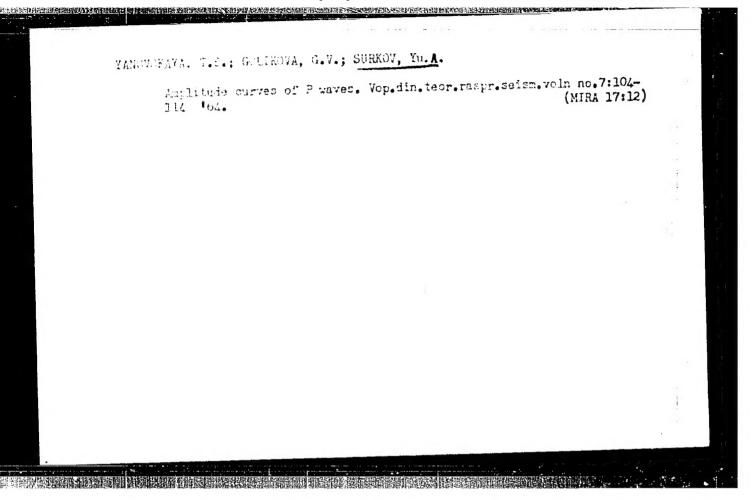
1. Vernadsky Institute of Geochemistry and Analytical Shemistry,
Academy of Sciences U.S.S.R., Moscow.

(Thallium—Analysis)



BARANOV, V.I.: BARSUKOV, V.L.; IVANOVA, V.F.; KERISTIANOV, V.K.; SUKKOV, Yu.A., kand. fiz.-matem. nauk, otv. red.

[Neutron methods of research and analysis of boroncontaining raw materials] Neitronnye metody poiskov i analiza bornogo syr'ia. [sy V.I.paranov i dr.] Moskva, Izd-vo "Nauka," 1964. 139 p. (MIRA 18:1)



TITLE: Measurement of Famma-radiation of the lunar surface of the Seventh COSPAR Meeting held in Vienna in May 1966]  SOURCE: Geokhimiya, no. 8, 1966, 891-899  TOPIC TAGS: radiation measurement, gamma radiation, moon, lunar probe, scintillation spectrometer  ABSTRACT: The spaceship Luna 10, placed into a selenocentric orbit on 3 April 1966, was equipped with a 32-channel scintillation spectrometer to investigate the intensity and spectral composition of y-radiation emitted from the lunar surface. The absence of an atmosphere sufficiently dense to absorb lunar surface. The absence of an atmosphere sufficiently dense to absorb lunar surface. The absence of an atmosphere sufficiently dense to absorb lunar surface. The absence of an atmosphere sufficiently dense to absorb lunar surface. The absence of an atmosphere sufficiently dense to absorb lunar surface. The absence of an atmosphere sufficiently dense to absorb lunar surface. The absence of an atmosphere sufficiently dense to absorb lunar surface. The absence of an atmosphere sufficiently dense to absorb lunar surface. The absence of an atmosphere sufficiently dense to absorb lunar surface. The absence of an atmosphere sufficiently dense to absorb lunar surface. The absence of an atmosphere sufficiently dense to absorb lunar surface. The absence of an atmosphere sufficiently dense to absorb lunar surface. The absence of an atmosphere sufficiently dense to absorb lunar surface. The absence of an atmosphere sufficiently dense to absorb lunar surface. The absence of an atmosphere sufficiently dense to absorb lunar surface. The absence of an atmosphere sufficiently dense to absorb lunar surface. The absence of an atmosphere sufficiently dense to absorb lunar surface in the solid angle subtended by the visible surface.	distribution of the second	L 04702-67 FSS-2/EWT(1)/EWT(m)/FCC JKT/TT/CW  ACC NR: AP6028010 SOURCE CODE: UR/0007/66/000/008/0891/0899  AUTHOR: Vinogradov, A. P.; Surkov, Yu. A.; Chernov, G. M.; Kirnozov, F. F.; & 2  Nazarkina, G. B.  ORG: Institute of Geochemistry and Analytical Chemistry im. V. I. Vernadskiy,  AN SSSR, Moscow (Institut geokhimii i analiticheskoy khimii AN SSSR)  AN SSSR, Moscow (Institut geokhimii i analiticheskoy khimii AN SSSR)  TITLE: Measurement of gamma-radiation of the lunar surface by the Luna-10 spaceship
1966, was equipped with a 52-chainter of γ-radiation emitted from the gate the intensity and spectral composition of γ-radiation emitted from the gate the intensity and spectral composition of γ-radiation dense to absorb lunar surface. The absence of an atmosphere sufficiently dense to absorb lunar surface. The absence of an atmosphere sufficiently dense to absorb lunar surface.		[Paper presented at the Seventh Goods, Source: Geokhimiya, no. 8, 1966, 891-899 TOPIC TAGS: radiation measurement, gamma radiation, moon, lunar probe,
		gate the intensity and spectral composition of $\gamma$ -radiation emitted from the gate the intensity and spectral composition of $\gamma$ -radiation emitted from the gate the intensity and spectral composition of $\gamma$ -radiation dense to absorb lunar surface. The absence of an atmosphere sufficiently dense to absorb lunar surface. The absence of an atmosphere sufficiently dense to absorb lunar surface.

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of the Moon, which was 0.9  $\pi$  at periselene and 0.46  $\pi$  at aposelene in the initial orbit.

\$P\$在这些是没有的现在分词,我们就会想到这种情况的。

It is known that the content of natural radioactive elements (U, Th,  $K^{40}$ ) in terrestrial rocks decreases from acidic to basic to ultrabasic rocks and that the decrease covers a range of several orders of magnitude. Therefore, it was expected that it would be possible to determine the type of rocks present in the lunar surface from the relative content of U, Th, and K established from the the ray spectrum. In conducting the experiments, the Y-radiation from natural radioactive elements can. be lower than the level of Fradiation produced during the interaction of primary cosmic particles (primarily protons) with the lunar surface was taken into account by analyzing the characteristic Frays emitted during the interaction.

Instrumentation

The measurements were made with a scintillation spectrometer consisting of a 3 x 4-cm NaI(Tl) cylindrical crystal Y-ray detector with an FEU-16 photomultiplier and a pulse-height analyzer. To eliminate the back-

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ground from charged particles, the NaI(Tl) crystal was enclosed in a container of a thin plastic scintillator. The pulses generated by charged particles were registered by the NaI(Tl) crystal and the plastic scintillator and were then separated from the pulses generated by \u03c4-rays which went practically unregistered by the plastic scintillator.

The scintillation spectrometer recorded Y-ray spectra in the energy ranges between 0.3-3-1 and 0.15-1.5 Mev. The switching of energy ranges was performed by ground command. The detector and the analyzer of the spectrometer were located in a hermetically sealed compartment under a shell 1 g/cm<sup>2</sup> thick.

#### Experimental Results

Six Y-ray spectra in the energy range 0.3—3.1 Mev were obtained during the first month of operation of Luna 10. In addition, the integrated intensity of Y-radiation in the same energy range was obtained at approximately points. The measurements were conducted over relatively wide surface areas covering the continents and the seas on both the light and the dark sides of the Moon. The height and the approximate selenographic coordinates

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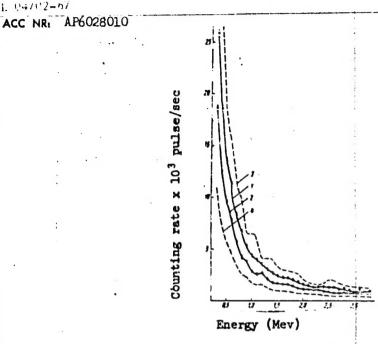


Fig. 1. Gamma-ray spectra obtained by Luna 10 while in orbit around the Moon and along the trajectory of the flight from the earth to the Moon

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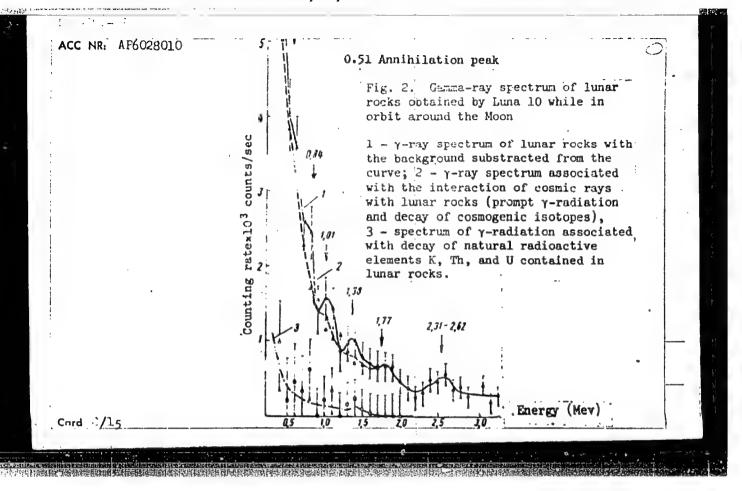
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 $1-\gamma$ -ray spectrum of the lunar rocks together with the background;  $2-\gamma$ -ray spectrum of the background due to interaction of cosmic rays with the material of Luna 10 corrected for the screening by the Moon; 3 and 4- same spectra as those given by 1 and 2, respectively, recalculated to represent measurements which would be taken at the surface of the Moon. The errors shown are root-mean-square errors.

interaction of cosmic rays with the substance of Luna 10, taking the screening by the moon into account, is also shown in Fig. 1 (curve 2).

Compared to the counting rate of rays measured along the flight trajectory, the counting rate in orbit around the Moon increased by 30-40%.

As a result of the screening effect of the Moon, the background due to irradiation of the spaceship by cosmic particles near the Moon decreases and is equal to about 78—89% of the background encountered along the trajectory of the flight. The background spectrum was measured during the flight



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hard Y-rays with energies in excess of 1.5 Mev while the spectrum of natural radioactive elements is characterized by a steep slope at higher energies and the absence of Y-rays with energies greater than 2.62 Mev. This shows that most Y-radiation from the lunar surface is not associated with the natural radioactivity of U, Th, and K<sup>40</sup> but is the result of the interaction of cosmic rays with the lunar substance and the decay of cosmogenic isotopes.

Table 2 shows the characteristic γ-rays identified from the lunar γ-ray spectra and the principal nuclear reactions involving the probable constituent elements of lunar rocks. It can be seen from Table 2 that O, Si, Al, and Mg are likely the most widely distributed elements in lunar rocks.

Table 2. Energies of Gamma Rays Identified From the Lunar Gamma-Ray Spectra

Energy (Mev)	Principal Nuclear Reactions Causing Emission of Characteristic Gamma-Rays
0,84 1,01 1,37 1,78 2,31 2,62	$\begin{array}{lll} \Lambda_1^{27}(p,p'\gamma) \; \Lambda_1^{27}, \; Si^{28}(p,2p\gamma) \; \Lambda_1^{27}, \; Fe^{84}(p,p'\gamma) \; Fe^{86} \\ \Lambda_1^{27}(p,p;\gamma) \; \Lambda_1^{28}, \; Si^{28}(p,2p\gamma) \; \Lambda_1^{127} \\ \Lambda_2^{24}(p,p'\gamma) \; Mg^{24}, \; \Lambda_1^{27}(p,pt\gamma) \; Mg^{24}, \; Si^{28}(p,px\gamma) \; Mg^{24} \\ \Lambda_2^{24}(p,p'\gamma) \; Ne^{28}, \; \Lambda_1^{27}(p,pt\gamma) \; Mg^{24}, \; Si^{28}(p,p'\gamma) \; Si^{28} \\ O^{14}(p,px\gamma) \; N^{14}, \; Mg^{34}(p,pn\gamma) \; Mg^{23}, \; Mg^{34}(p,2p\gamma) \; Na^{23}, \; Al^{27}(p,pln\gamma) \; Mg^{23} \end{array}$

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Since the lunar surface is exposed to constant interaction with cosmic rays, all of the cosmogenic radioisotopes should be in radioactive equilibrium. Therefore, both long-lived and short-lived radioisotopes should be radioactive, and their content should be proportional to the effective cross section for their production. Calculations show that the main contribution to Y-ray emission is made by the decay of the following cosmogenic isotopes:  $O^{14}(T_{1/2} = 72 \text{ sec}, E_{\gamma} = 2.31 \text{ Mev}), O^{19}(T_{1/2} = 27 \text{ sec}, E_{\gamma} = 1.37 \text{ Mev}), F^{20}(T_{1/2} = 10.7 \text{ sec}, E_{\gamma} = 1.63 \text{ Mev}), Na^{22}(T_{1/2} = 2.6 \text{ hr}, E_{\gamma} = 1.28 \text{ Mev}), Na^{24}(T_{1/2} = 15 \text{ hr}, E_{\gamma} = 1.37 \text{ Mev} \text{ and } 2.76 \text{ Mev}).$  These radioisotopes are formed with a considerable yield in nuclear reactions involving the same rock-forming elements: Mg, Al, and Si.

The peak at 0.51 Mev, which is especially pronounced in the lunar Y-ray spectra measured in the energy range 0.15—1.5 Mev, is produced by Y-radiation emitted during annihilation.

Analysis of the results shows that the Y-radiation intensity corrected for the difference in height is practically constant above the different regions of the lunar surface (intensities did not differ by more than 40%). This can probably be attributed to the fact that the main source of Y-rays is cosmic radiation. A preliminary analysis shows that the total dose rate of

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assumption follows from the fact that the spaceship and its components were made of light alloys of Si, Al, and Mg with very little Fe, i. e., the dominant elements in the composition of rocks. Curve 3 in Fig. 2, obtained by subtracting curve 2 from curve 1, shows the Y-ray spectrum produced by the decay of natural radioactive elements. Fig. 2 shows that 90% of the intensity of gamma radiation emitted by lumar rocks is produced by radioactivity induced by cosmic rays and no more than 10% by decay of K, Th, and U.

Prior to the flight the Y-spectrometer aboard the spaceship was precalibrated using samples with a measured amount of K, Th, and U and also with rock samples containing different amounts of these elements. This procedure made it possible to calculate the Y-ray spectra, which should be obtained by the orbiting spaceship, emitted by rocks with different amounts of natural radioactive elements (it was assumed that the radiation produced by cosmic rays is absent). Fig. 3 shows such spectra which would be obtained at a height of 350 km with the background subtracted from the spectrum. The hatched areas correspond to range of concentrations of radioactive elements for given types of rock. The average values of concentrations of K, Th, and U were taken from a paper by A. P. Vinogradov (Geokhimiya, no. 7, 1962).

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1. 04762-67 ACC NR: AP6028010

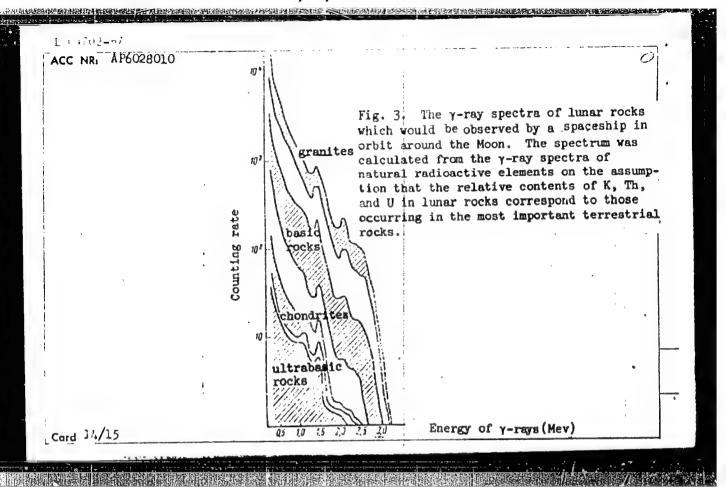
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Comparison of the lunar y-ray spectra with those of terrestrial rocks with a known content of K, Th, and U shows that at least in the regions of the Moon over which measurements were conducted there are no rocks on the lunar surface, or at a depth not exceeding 27 cm, containing the same amount of K, Th, and U as do the acidic terrestrial rocks, such as granites. The intensity of \gamma-radiation due to natural radioactivity (Fig. 2, curve 3) tends to indicate the presence of basic rocks such as basalts. However, at the present time it is impossible to exclude the possibility that the concentration of natural radioactive elements was estimated a bit too high. It is interesting to note that tektites, which have almost the same composition and amounts of U, Th, and K as acidic rocks, cannot be of lunar origin.

#### Conclusions

- The main results obtained from the measurements of the intensity and spectral composition of radiation by the Luna 10 can be summarized al follows:
- 1. The overall level of Y-radiation of the lunar surface slightly exceeds that of the Earth. Preliminary results show that the intensity of Y-radiation of

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L Children ACC NR: AP6028010 the surface of the Moon is  $20-30 \mu r/h$ . 2. About 90% of the Y-rays emitted by the surface of the Moon is produced during interaction of cosmic rays with lunar substance (prompt Y-rays and decays of cosmogenic isotopes). 3. The basic rock-forming elements of the lunar surface are believed to be 0, Mg, Al, and Si. 4. No difference was noted in intensity of γ-rays emitted by different regions of the lunar surface including the seas and the continents (variation of intensity did not exceed 40%). 5. The decay of K, Th, and U in lunar rocks does not contribute more than 10% to the total \u03c4-ray emission of the lunar surface. 6. Comparison of the intensity of \u03c4-radiation from the decay of natural radioactive elements K, Th, and U with the results obtained by a calibrated instrument from terrestrial rocks shows that the concentration of radioactive elements in lunar rocks is close to that of basic terrestrial rocks and differs greatly from acidic rocks. However, it can not be positively stated that the lunar surface contains no ultrabasic (meteoritic) substance. At the present time an attempt is being made to determine the relative content of 0, Mg, Al, and Si in lunar rocks from the available Y-ray spectra produced in interactions with cosmic rays. Orig. art. has: 3 figures and 3 tables. [FSB: v. 2, no. 10] SUB CODE: / SUBM DATE: 24Jun66 / ORIG REF: 002 Card 15/15

L 14960-66 EWT(1)/EWT(m)/EWA(d) GW

ACC NR: AP5022952

SOURCE CODE: UR/0007/65/000/008/0918/093

AUTHOR: Surkov, Yu. A.; Nazarkina, G. B.

ORG: Institute of Geochemistry and Analytical Chemistry im. V. I. Vernadskiy.

SSSR, Moscow (Institut geokhimii i analiticheskoy khimii AN SSSR)

TITLE: Nuclear reactions in meteorites

SOURCE: Geokhimiya, no. 8, 1965, 918-935

TOPIC TAGS: meteorite, radioactive isotope, nuclear reaction, cosmic ray effect,

astrophysics, cosmogony

ABSTRACT: In a brief survey of the literature, the authors discuss various problems which have recently arisen with the analysis of isotopes formed by cosmic rays in meteorites and other objects from outer space. The sources considered in this paper are divided into two categories: astrophysical and nuclear. The first deals with problems in the origin and evolution of meteorites and cosmic rays, while the second is concerned with a study of the structure of atomic nuclei and the mechanism of nuclear reactions under the influence of high energy particles. Tables are given showing the radioactivity of cosmogonic isotopes as well as the concentration of

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stable isotopes in iron and stony meteorites. The cosmic age of meteorites is determined from the ratio of the concentration of cosmogonic stable nuclei to their rate of formation at the present time. It is found that most of the stony meteorites have a cosmic age of  $10^6 - 6 \cdot 10^7$  years. The cosmic age of iron meteorites is found to be 1-7-108 years. All meteorites whose ages have been measured are younger than the geological age of the earth, which is ~4.5·109 years. It is concluded that all meteorites were produced in one or two gigantic catastrophies. Curves are given showing the accumulation of cosmogonic isotopes in a meteorite. These curves show that the number of stable isotopes increases constantly. Shoft bursts of cosmic rays have practically no effect on the radioactivity of long-lived isotopes. Short-lived isotopes quickly reach equilibrium and the curve for their radioactivity with respect to time is practically a repetition of that for the intensity of cosmic rays. After the meteorite falls to the earth, the radioactivity of the corresponding isotopes falls exponentially. The number of stable isotopes continues to increase even after the fall. The integral spectrum of protons and alpha particles from primary cosmic radiation is given. Rocket data show that the integral intensity of cosmic particles of galatic origin at great distances from the earth is √2.5 particles/cm<sup>2</sup>. A curve is given showing the integral spectrum of cosmic rays

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generated during solar flares. Hypotheses explaining the mechanisms responsible for various nuclear phenomena caused by high energy particles are discussed. Experimental data indicate that cosmogonic isotopes are formed in different meteorites under identical conditions. However, the differences in the meteorites with respect to composition and size as well as the difference in ages cause wide variation in the concentration of cosmogonic isotopes. Orig. art. has: 5 figures, 2 tables, 10 formulas.

SUB CODE: 03,04 / SUBM DATE: 00/ ORIG REF: 024/ OTH REF: 111

Card 3/3 AC

ACC NR: AP7005449

SOURCE CODE: UR/0020/66/170/003/0561/0564

AUTHOR: Vinogradov, A. P. (Academician); Surkov, Yu. A.; Chernov, G. M.

O.G: Institute of Goochemistry and Analytical Chemistry in. V. I. Vernadskiy, AN SSSR (Institut gookhimii i analitichoskoy khimii AN SSSR)

TIPLE: Investigations of the intensity and spectral composition of lunar gasma radiation on the automatic station "Luna-10"

SOURCE: AM SSSR. Doklady, v. 170, no. 3, 1966, 561-564

TOPIC TAGS: gamma spectrum, gamma spectrometer, scintillation spectrometer, bromsstrahlung, cosmic radiation, moson, lunar satellite, photomultiplier/Luna-10 lunar satellite, FEU-16 photomultiplier

ABSTRACT: "Luna-10" carried a scintillation gamma spectrometer with a detector of Y-radiation; this was a NaI(T1) crystal measuring 30 x 40 mm, connected to a FAU-16... photocultiplier, and a pulse amplitude analyzor. The instrument made it possible to measure the spectrum of Y-radiation against a background of charged particles. The instrument recorded the spectrum of Y-radiation in two ranges: from 0.3 to 3.1 MeV and from 0.15 to 1.5 MeV. During the first menth of operation of "Luna-10" it was possible to obtain 6 spectra of Y-radiation in the energy range from 0.3 to 3.1 MeV. In addition, at approximately 15 points the intensity of Y-radiation was measured in this same range of energies. The measurements covered rather extensive areas of the surface, including both the "continents" and "seas" on both the visible and far sides. Analysis of the form of the lunar Y-spectra revealed that they differ

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ACC NS: A.7007599 SOURCE CODE: UR/0293/66/004/006/0871/0879

AUTHOR: Vinogradov, A. P.; Surkov, Yu. A.; Chernov, G. M.; Kirnozov, F. F.; TITLE: Preliminary results of measurements of gamma radiation of the lunar surface on the space station "Luna-10" SOURCE: Kosmicheskiye issledovaniya, v. 4, no. 6, 1966, 871-879 TOPIC TAGS: lunar satellite, gamma spectrometer, cosmic radiation SUB CODE: 22, 20,18 ABSTRACT: Experimental investigations of the intensity and spectral componsition of gamma radiation of lunar rocks made using a gamma spectrometer carried aboard the automatic station "Luna-10" demonstrated that: 1) The general level of gamma radiation of lunar rocks approaches the level of gamma radiation over rocks of the earth's crust, somewhat exceeding the latter. According to a preliminary estimate, the intensity of the gamma radiation at the lunar surface is 20-30 uR/hour. 2) The principal contribution to lunar gamma radiation is from processes of the interaction of cosmic rays with lunar matter (instantaneous gamma radiation and the decay of cosmogenic isotopes). About 90% of the total lunar gamma radiation can be attributed to these processes. 3) Analysis made it possible to identify in the lunar spectrum photopeaks from gamma quanta emitted during the interaction of cosmic particles with the principal rock-forming elements of the lunar surface -- 0, Mg. Al, Si -- and gamma quanta emitted during the decay of cosmogenic isotopes. 4) The results of measurements over different regions of the lunar surfaces, including the regions of the lunar "continents" and Seas" did not make 629.195.3:523.36 UDC:

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ACC NR. AP7005118 SOURCE CODE: UR/0007/66/000/008/0891/0899 AUTHOR: Vinogradov, A. P.; Surkov, Yu. A.; Chernov, G. M.; Kirnozov, F. F.; Nazarkina, G. B. ORG: Institute of Geochemistry and Analytical Chemistry im. V. I. Vormadskiy, AN SSSR. Poscow (Institut geokhimii i analiticheskoy khimii AN SSSR) TITLE: Measurements of gamma radiation of the lunar surface by the space station Luna-lo SOURCE: Geokhimiya, no. 8, 1966, 891-899 TOPIC TAGS: gamma spectrum, lunar satellite, earth crust, lunar surface, lunar' radiation / Luna-10 lunar satellite ABSTRACT: During its first month of operation the lunar satellite "Luna-10" obtained six spectra of gamma radiation in the energy range from 0.3 to 3.1 MeV. In addition, at about 15 points it measured the total intensity of gamma radiation in the same energy range. The measurements covered extensive areas of the surface of both the seas and continents on both sides of the moon. It was found that the general level of gamma radiation of lunar rocks approaches the level of gamma radiation over ... the rocks of the earth's crust, somewhat exceeding the latter. The preliminary estimate of gamma radiation for the lunar surface is 20-30 µ curies. The principal contribution to lunar gamma radiation is from processes of interaction of cosmic rays with lunar matter (instantaneous gamma radiation and the decay of isotopes). About 90% of the total lunar gamma radiation can be attributed to these processes. Analysis

CC NR: AP7005118

makes it possible to identify in the lunar spectrum photopeaks from gamma quanta cmitted at the time of interaction between cosmic particles and the principal rock-forming elements of the lunar surface -O. Mg. Al. Si. as well as gamma quanta emitted during the decay of cosmogenic isotopes. (The possibility of determining the relative content of these elements now is being studied.) Results of measurements over different regions of the lunar surface, including the seas and continents, did not reveal an appreciable difference in the intensity of gamma radiation over these regions (intensity variations do not ; exceed 40%). In the total intensity of gamma radiation of lunar rocks the percentage of radiation caused by decay of K, Th and U is approximately 10%. Comparison of the intensity of gamma radiation from decay of the natural radioactive elements K, Th and U with the results of calibration of the instrument against terrestrial rocks makes it possible to ascribe to lunar rocks concentrations of radioactive elements close to terrestrial rocks of basic composition (such as basalts). The data indicate that there are no areas of rocks with concentrations of radivactive elements such as terrestrial granites, and especially none With ore concentrations of K, Th and U. Orig. art. has: 3 figures and 3 tables. /JPRS: 38,460/

SUB CODE: 03,22,20 / SUBM DATE: 24Jun66 / ORIG REF: 002

ard 2/2

S/124/61/000/009/035/058 D234/D303

AUTHOR:

Surkov, Yu.P.

TITLE:

Effect of the clasticity of fixing ends on the

natural vibrations of rods

PERIODICAL:

Referativnyy zhurnal. Hekhanika, no. 9, 1961, 13,

abstract 9 V100 (Uch. zap. Kuybyshevsk. gos. ped.

in-t, 1959, no. 29, 79-84)

TEXT: Longitudinal and transverse vibrations of rods with clastically fixed ends are considered. The problem is solved by the method of perturbations. The parameters which enter into the boundary conditions are regarded as small which allows one to look for the eigenfunctions and eigenvalues of the problem in the form of series expansion in powers of these parameters. The solution of the undisturbed problem, differing from the former by the parameters of the boundary conditions being equal to 0, is assumed to be known. "Additions" to the solution of the undisturbed problem

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ZUBUV, V. Ya., doktor tekhn.rauk, prof.; SOKDLOV, N.V., kand.tekhn.nauk;

KHIGIL NIKOV, L.A., inzh.; ERACHEV, S. V. inzh.; Prinimali uchastiye:

Deformation of metastable austenite and the strength of steel
strip. Stal 21 no.6:549-551 Je '61.

1. Beloretskiy staleprovolochnyy zavod i Ural 'skiy politekhnicheskiy institut.

(Steel.-Hardening)

(Metallography)

Card 1/3

只能在这种的**是这种特别的证明,因为 化基内线图式**的图象的 经加速机构的图像的 上海市 的复数形式 经对外收益 医多次氏性 化二甲基甲基甲基甲基甲基甲基甲基甲基甲基甲基甲基 33461 5/129/62/000/001/004/011 1754 1042 E073/E483 1 1700 Zubov, V.Ya., Doctor of Technical Sciences, AUTHORS: Grather, S.V., Surkov, Yu.P. Engineers Influence of thermomechanical treatment on the TITLE . strength of steel wire PERIODICAL: Metallovedeniye i termicheskaya obrabotka metallov, no.1, 1962, 20-22 The authors studied the possibility of using thermomechanical treatment in the drawing of carbon- and low-allcy steel wire (L.A.Krasil nikov and A.G.Lysenke participated in the The chemical compositions (%) of the steels experiments), investigated are given as follows: Cu Na Sí Cr Mn C 0.025 0.020 0.10 0.11 80.0 0.21 0.27 0.71 √ 7 A (U7A) 0.006 0.019 0.20 0,12 0.18 51.0 1.01 0.20 710Å (U10A) 0.023 0.20 0.10 0.019 0.09 0.98 0.23 0.66 651" (65G) 0,009 0.017 0,10 DM 142 (E1142) 0.66 0.50 1.67 0.33 The wire was The initial wire diameters were 1.95 and 2 mm, heated for the drawing operation to 920 to 940°C by passing an

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的一个人,我们就是一个人的人,我们就是一个人的人,我们就是一个人的人,我们就会会会会会会会会会。

Influence of thermomechanical ...

33461 \$/129/62/000/001/004/011 E073/E483

electric current through it and cooled in a lead bath to 320 - 350 C. The speed of movement of the wire was 10 m/min. The wire was deformed in a single pass (short incubation period) by 5 to 32%, using a soap-graphite lubricant. The final cooling after drawing was in air. After this, thermomechanical treatment specimens of the wire were tempered under laboratory conditions at 100 to 500 C with a holding time of 1.5 min. strength of the wire drawn whilst the austenite was in the super cooled state was very high. It was highest for the steel EII43. 1,e, 306 kg/mm2 (32% reduction and tempering at 350°C for 3 min). Further experiments were carried out exclusively on this material. The hardness after thermomechanical treatment was higher by about a HRC units than for the same material quenched in the ordinary way. In addition, hardness of thermomethanicallytreated material decreased more slowly with intreasing tempering temperature than that of the same material after step-wise quenching. These differences were attributed to smaller grainsize of marrensite and presence of strain hardened austenite in thermomethanically treated steel. Even after Card 233

33461

Influence of thermomechanical ...

S/129/62/000/001/004/017 E073/E483

3 minutes tempering at relatively high temperatures (450 to 500°C), hardness of thermomechanically treated steel was practically equal to that of the same material after the conventional hardening reduction but there was a drop in strength after reductions not exceeding 6 to 8%. If reductions of the order of 30% are used (followed by tempering for 1 min at temperatures not exceeding with a strength of the order of 300 kg/mm<sup>2</sup>. There are 5 figures,

ASSOCIATION: Ural skiy politekhnicheskiy institut
(Ural Polytechnical Institute)

Card 3/3

S/0126/63/016/006/0918/0921

ACCESSION NO: AP4009386

AUTHORS: Sokolkov, Ye, N.; Surkov, Yu. P.

TITLE: Substructure formation in austenite steel during plastic deformation at high temperatures

SOURCE: Fizika metallov i metallovedeniye, v. 16, no. 6, 1963, 918-921

TOPIC TAGS: austenite steel, steel, steel substructure, plastic deformation, high temperature deformation, Cr-Ni-Mn steel, W alloy, Mo alloy, steel microstructure

ABSTRACT: The effect of temperature and deformation velocity on the formation of substructure in steel was studied in the process of plastic deformation at high temperatures. The samples consisted of Cr-Mn austenite steel (0.37% C) alloyed with W and Mo. Their substructure (revealed by etching) was studied microscopically and the structure patterns were photographed. It was established that a plastic deformation of the metal at 1000-1200C and at deformation velocity 0.015-0.0035 1/sec caused the appearance of substructure. The size of the substructure elements increased with the increase in temperature and with the decrease in deformation speed. The origin of the substructure observed was explained by the redistribution of dislocations in the process of deformation at high temperatures.

**Card** 1/2

ACCESSION NO: AP4009386

"We express our gratitude to V. D. Sadovskiy for his participation in the discussion of the results. V. A. Yudin conducted the electron-microscope studies."

Orig. art. has: 2 figures.

ASSOCIATION: Institut fiziki metallov AN SSSR (Institute of Physical Metallurgy AN SSSR)

SUBMITTED: 28Jul63

DATE ACQ: 03Feb64

ENCL: 00

SUB CODE: ML

NO REF SOV: 005

OTHER: 006

Card 2/2

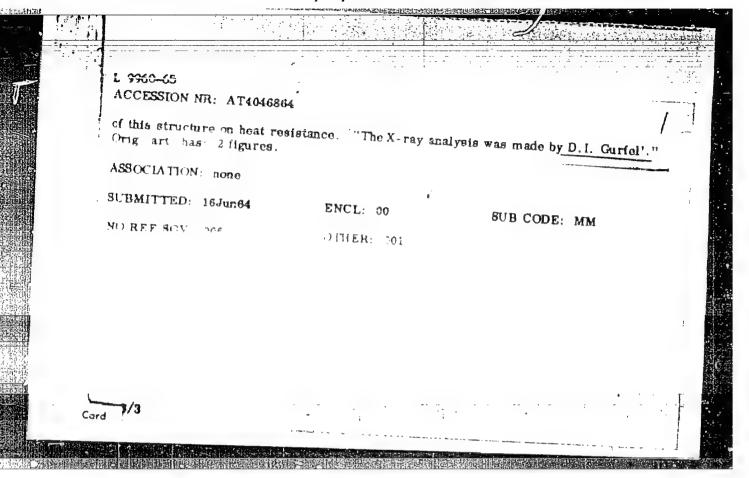
THE ATTEMPT OF THE STATE OF A DOMESTIC AUDITOR TO THE MEN 8/ 0000/ 64/ 000/ 000/ 0331/ 0335 ACCESSION NR: AT4046864 AUTHOR: Sokolkov, Ye. N., Sadovskiy, V.D., Surkov, Yu. P., Chuprakovs, N.P. Nichkova, M.M. TITLE: Investigation of the hardening and structural stability of austenitic alloys after tion tomperature thermomechanical treatment SCURCE: AN SSSR. Nauchny'y sovet po probleme zharoprochny'y splavov. Issledovaniya the cases and manufactured treatment allow hardening, allow structure, allow that the second of the second Austra Colonia is the remember of next resistance by high-temperature the commentanical the control of appearal structure in the material during bot and the state of t numberatic steel with admixtures of tungsten and titanium after high-temperature thermome hapte is treatment and aging. The effect of temperation will plastic deformation rate

1. 9960-65

ACCESSION NR: AT4046864

were studied in relation to recrystallization in alloys rolled at a rate of 1.5 m/min at 350-1150C. Samplesfor rolling were 11.5x11.5x60 mm, and for upsetting were 50x50x70 mm. Special insulation was used during upsetting to prevent rapid cooling. All samples were cooled in water after plastic deformation. The effects of aging were studied by hardness measurements, while structural stability was measured by microstructural analysis. Hardness measurements showed that all alloys selected in the test showed a higher hardness than after the usual thermal treatment. The difference in hardness in comparis a with the usual hardening procedures was 15-20 kg/mm² even after high temperature thermomechanical treatment at 800C for 32 hours. Similar results were about ed for other heating and aging temperatures. Impact toughness was also higher after high temperature thermomechanical treatment (12-13 kg-m cm² instead of 5-7).

avoiding formulation of the fisher analysis also showed in a lowering of the deformation rate to reased the tendency toward recrystallization, and altered the grain boundary deviations from a dentate to a wave-like pattern. These results lead to new possibilities for applying high-temperature thermomechanical in atment more attraction. Special investigations will be required, however, to find the effect



L 8914-66 EWT (m) /EHA (d) /EWP (t) /EWP (z) /EWP (b) UTW/JD

ACC NR: AP5027141 UR/0126/6

UR/0126/65/020/004/0561/0565

AUTHOR: Sokolkov, Ye. N.; Surkov, Yu. P., Gurfel', D. I.

ORG: Institute for the Physics of Metals, AN SSSR (Institut fiziki metallov AN SSSR)

TITLE: Effect of conditions of high temperature heat and mechanical treatment on the thin crystalline structure of chromium-mickel-manganese austenitic steel

SOURCE: Fizika metallov i metallovedeniye, v. 20, no. 4, 1965, ... 561-565

TOPIC TAGS: crystal structure, austenite steel, chromium steel, nickel steel, manganese steel, work hardening, metal heat treatment, plastic deformation / EI481 chromium steel /

ABSTRACT: A study was made of the fine crystal structure of chromium-nickel-manganese steel <u>EI481</u>! as a function of the conditions of high temperature heat and mechanical treatment: temperature and degree and rate of deformation. Samples with dimensions 50 x 50 x 75 mm were heated to 1200°C, held there for 1.5 hours, and deformed at this temperature and at 1100 and 1000° after cooling in the furnace

Card 1/3

UDC: 669.15.018.45

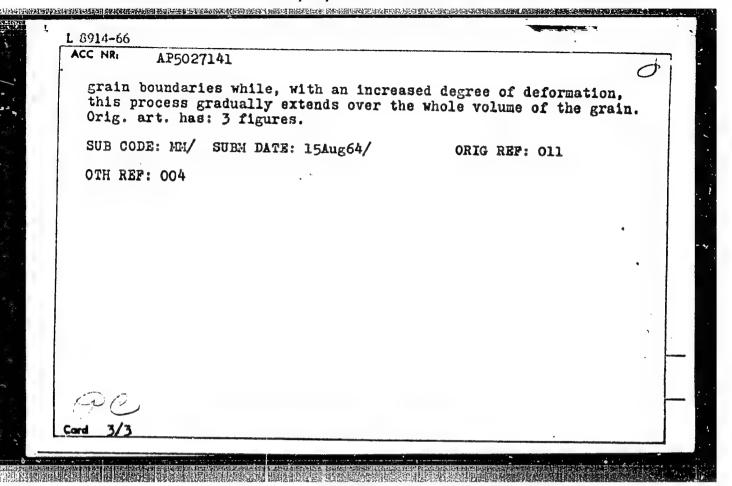
L 8914-66

ACC NRI AP5027141

Deformation was done by 25-30% upsetting in a press at a rate of 0.015, 0.0075, and 0.005 1/sec (the degree of deformation was evaluated by the change in height of the sample). After deformation (within 1-2 seconds), the samples were quenched in water. The test samples were subjected to metallographic and x-ray structural analysis. of the microstructure (illustrated in the article) shows that, as a result of high temperature heat and mechanical treatment, there appears a structure whose elements depend substantially on temperature and the rate and degree of plastic deformation. Treatment at 12000 at a minimum deformation rate leads to formation of subgrains with an average size of 30-40 microns. A decrease in deformation temperature to 1100 decreases the size of the subgrains to 15-20 microns. X-ray studies show that, in samples which have undergone conventional annealing, the grains have a sufficiently clear character with a small radial washing out, which probably indicates a certain elastic microdeformation of the lattice. For material subjected to high temperature heat and mechanical treatment, the x-ray studies indicate the formation within the grains of large mutually unoriented regions of the crystal lattice, that is, fragments. The magnitude of the plastic deformation has a complicated effect on the formation of the thin crystalline structure. At small reductions, the fragmentation of the structure is observed mainly in regions near the

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cm. 2/3



ACC NR: AT6034463

(A)

SOURCE CODE: UR/0000/66/000/000/0265/0271

AUTHOR: . Surkov, Yu. P.; Sadovskiy, V. D.; Sokolkov, Ye. N.; Pavlov, V. A.; Gaydukov, M. G.

ORG: none

TITIE: Effect of high temperature thermomechanical treatment at a small deformation rate on the heat resistance of Type KhN77YuR alloy

SOURCE: AN SSSR. Institut metallurgii. Svoystva i primeneniye zharoprochnykh splavov (Properties and application of heat resistant alloys). Moscow, Izd-vo Nauka, 1966, 265-271

TOPIC TAGS: metal heat treatment, heat resistance, metal deformation, metal recrystallization

ABSTRACT: High temperature thermomechanical treatment, concluding with deformation of the material at increased temperatures, and then cooling, eliminating the development of recrystallization due to the birth and growth of new grains, leads to a considerable improvement in the heat resistance properties of steels and alloys. The present article considers the effect of high temperature thermomechanical treatment at a small deformation rate (0.003-0.004 sec-1) on the heat resistance of alloy KhN77TYuR. Samples with a size of 50 x 50 x 75 mm were heated to a temperature of

Card 1/2

### ACC NRI AT6034463

1080° with a holding time of 8 hours, after which part of them were cooled in air (control samples), while the other part was subjected at the same temperature to 25-30% deformation. On the basis of the experimental results, the following conclusions were drawn: 1) high temperature thermomechanical treatment of alloy KhN77TYuR with a deformation rate of 0.003-0.005 sec<sup>-1</sup> assures a recrystallization structure in a cross section of the order of 50 x 50 mm, and leads to an improvement in heat resistance properties; 2) fragmentation of the structure in the alloy assures greater stability, and increases the temperature of articles made from the alloy by the method of high temperature thermomechanical treatment (up to 850°). Orig. art. has: 3 figures and 1 table.

SUB CODE: 11/ SUBM DATE: 10Jun66/ ORIG REF: 005

Card 2/2

RABINOVICH, V.S.; SURKOV, V.D.; SURKOVA, A.A.

Giardiasis in children. Pediatrila 37 no.7:88 J1 '59.

(MIRA 12:10)

1. Iz detskogo otdeleniya bol'nitsy imeni N.A.Senashko g.Iaro-elavlya.

(GIARDIASIS)

KOSEKO, i.e. eTIATUV 8.5.: 3788.07A, A.E.

Air mulling for seismie prospecting, Radved, 1 okt. metr.

OTERA 18-4.\*

1. Miskevskiy emsena Trudovego Kradnego Zameni institut
metrekhimishenkey 1 gazovey promysblennesti imeni skadenika
J.M.Gubkina.

### SURKOVA, A. V.; BLOSHTKYN, Ye. A.

Making forging die impressions by electric pulse techniques. Avt. prom. 29 no.5:38-39 My '63. (MIRA 16:4)

1. Mauchno-issledovatel skiy institut avtomobil noy promyshlennosti.

(Electric metal cutting)

OSTROVSKAYA, Sh. M.; YASINSKIY, A.V.; SURKOVA, D.F.

Materials on the epidemiology of Q fever in Tajikistan. Zdrav. Tadzh. 3 no.2:27-30 Mr-Ap 156 (MIRA 12:7)

1. Iz Stalinabadskogo Instituta epidemiologii i gigiyeny (dir. -dotsent M.Ya. Rasulov).

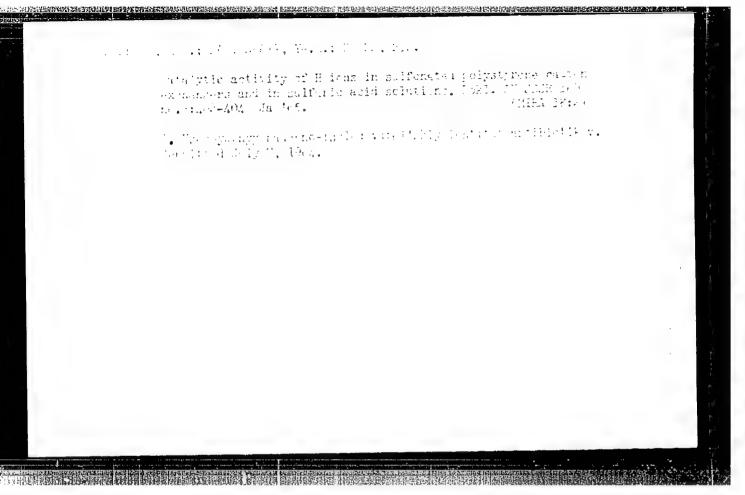
(TAJIKISTAN--Q FEVER)

OSTROVSKAYA, Sh.M.; YASHESKIY, A.V.; SURKOVA, D.F.

Results of a four-year study of Q fever in Yajikistan from 1953 to 1956. Zdrav.Tadsh. 6 no.4:18-22 J1-Ag '59. (MIRA 12:11)

1. Iz Stalinabadskogo instituta epidemiologii i gigiyeny.

(TAJIKISTAN--Q FEVER)

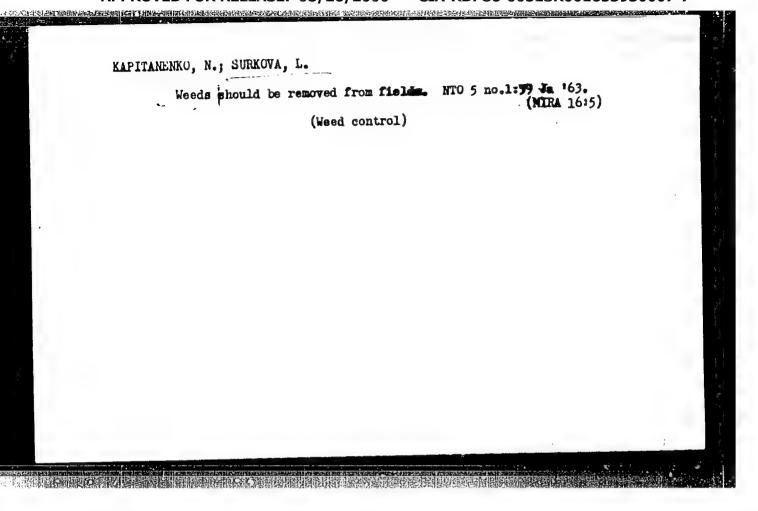


#### SURKOVA, L.

This should be the concern of all members of the plant. Pozh. delo 6 no.1:6 Ja '60. (MIRA 13:5)

1. Glavnyy inzhenor Moskovskogo khimiko-farmatsevticheskogo zavoda imeni N.A.Semashko.

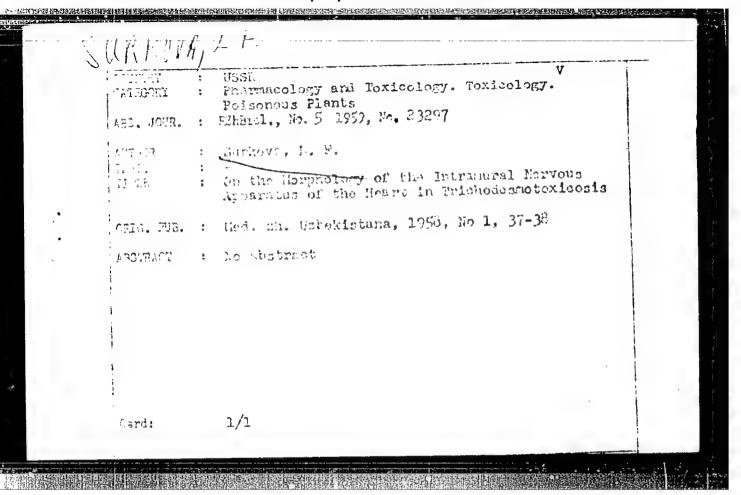
(Chemical plants-Fires and fire prevention)



LAM IN, Serafim Ivanovich, kand. veter. nauk; SURKOVA, L.A., red.

[Use of macro- and microelements in endemic diseases of animals] Primenenie makro- i mikroelementov pri endemicheskikh zabolevaniiakh zhivotnykh. Ulan-Ude, Buriatskoe knizhnoe izd-vo, 1964. 34 p. (MIRA 18:2)

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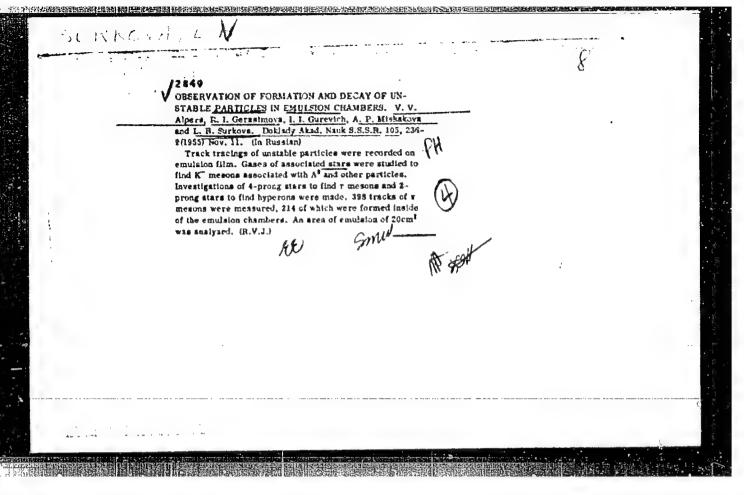


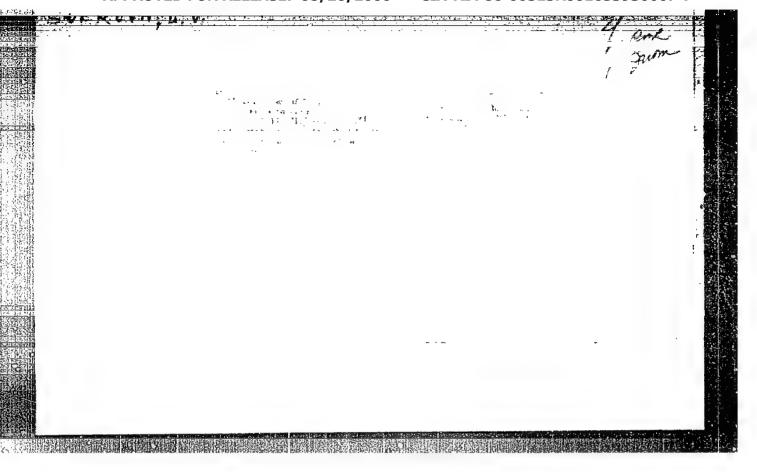
TABARTOVA, F.S., SUPPOVA, L.I.

Determining the bremine numbers of organosilison compounds containing unsaturated radicals. Plast. massy no. 12:51-52 165 (MEA 19:1)

#### "APPROVED FOR RELEASE: 08/26/2000

CIA-RDP86-00513R001653930007-7

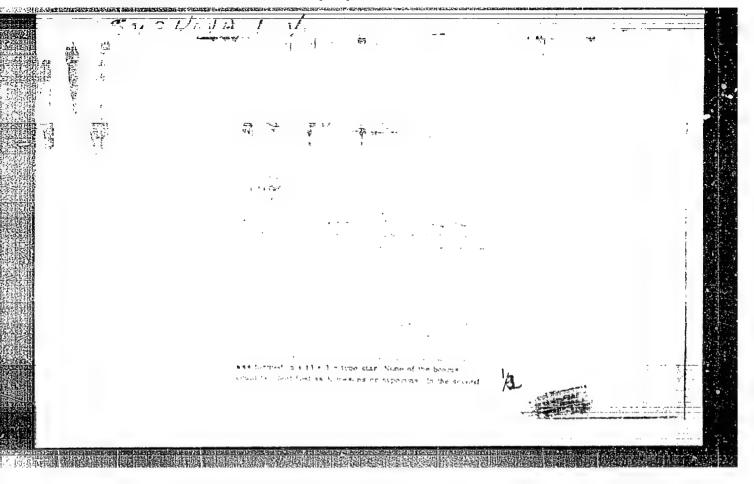


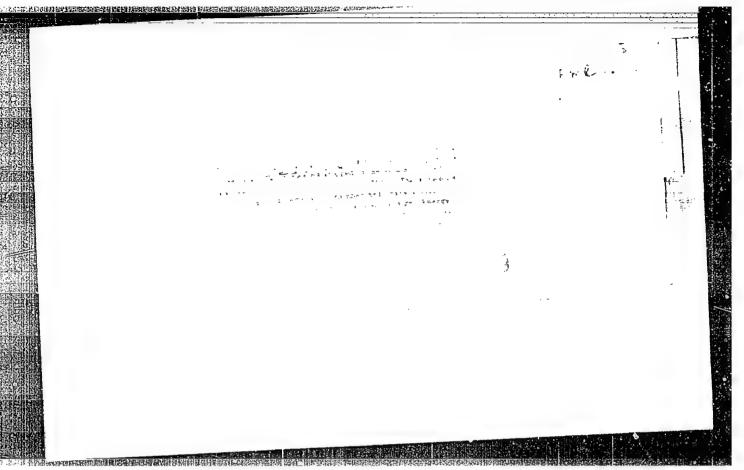


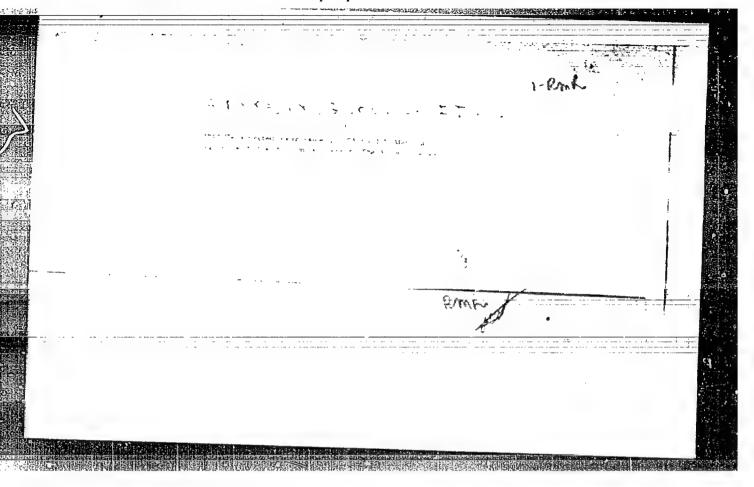
ALPERS, V.V.; GUREVICH, I.I.; SURKOVA, L.V.

Emulsion chamber observations on unstable particle production and decay. Dokl.AN SSSR no.3:421-422 My '56. (MLRA 9:8)

 Predstavleno akademikom L.A. Artsimovichem.
 (Cosmic rays) (Ionization chambers) (Photography, (Photography, Particle track)







PA - 2046

AUTHOR:

AL'PERS, V.V., GUREVIC, I.I., KUTUKOVA, V.M., MISAKOVA, A.P.

NIKOL'SKIJ, B.A., SURKOVA, L.V.

TITLE:

The Study of Explosion Showers produced by High Energy

Cosmic Particles (Russian).

PERIODICAL:

Doklady Akademii Nauk SSSR, 1957, Vol 112, Nr 1, pp 33-36

(U.S.S.R.)

Received: 2 / 1957

Reviewed: 3 / 1957

ABSTRACT:

The present work deals with the preliminary results obtained by studying 29 showers by the method of the emulsion chamber. The emulsion chamber consisted of 100 layers of 10 cm diameter and 450 & thickness. This emulsion chamber was irradiated in May 1955 for 7 hours at a height of 27 km. On the occasion of the microscopic investigation of these emulsions the explosion showers were fixed with more than 5 relativistic traces which are in a sufficiently narrow cone. Furthermore, the rays were fixed with more than 3 relativistic traces. On the occasion of the examination of 26,5 cm<sup>3</sup> photoemulsion 27 explosion showers and 29 rays were found. In the course of a further investigation of the rays through the emulsion chamber it was found that two of them originated from stars. The remaining 27 rays were found to be electron-photon showers. On the occasion of the microscopic investigation of the explosion showers the primary particle which excites the shower, the

Card 1/3

PA - 2046

The Study of Explosion Showers produced by High Energy Cosmic Particles (Russian).

are obtained on this occasion, are explicitly given for the case of ultrarelativistic shower particles. By assuming a nucleon-nucleon production mechanism of the shower we find  $n_s = k\sqrt{\cot n_1/2}$ . Some showers satisfy this relation and can thus be assigned to nucleon-nucleon interaction. However, the angular distributions of the shower particles contradict this conclusion, for a noticeable asymmetry of angular distribution was found. All showers produced by nucleons and  $\alpha$ -particles have a marked asymmetry with respect to the angle  $\theta = \pi/2$ .

ASSOCIATION: Mot given

PRESENTED BY: SUBMITTED:

AVAILABLE: Library of Congress

Card 3/3

GUREVICH, I.I.; MISHAKOVA, A.P.; NIKOL'SKIY, B.A.; SURKOVA, L.V.

Explosion showers produced by high energy cosmic ray particles. Zhur. eksp., i teor. fiz. 34 no.2:265-280 Y '59. (MIRA 11:4)

1. Akademiya nauk SSSR. (Cosmic rays)

Junkowa L.V.

AUTHORS:

Gurevich, I. I., Kutukova, V. H., Lishakova, 56-2-2/51

A. P., Nikol'shiy, B. A., Surkova, L. V.

TITLE:

The Asymmetry in the Angular Distribution of  $\mu^+$   $\rightarrow$  e<sup>+</sup>

Decay Electrons Observed in Photographic Emulsions

(Asimmetriya uglovogo raspredeleniya elektronov μ+ -> e+

-raspada po nablyudeniyam v fetoemul'sii)

PERIODICAL:

Churnal Eksperimental'noy i Teoreticheskoy Fiziki, 1958,

Vol 34, Nr 2, pp 280-285 (USSR)

ABSTRACT:

An emulsion chamber of 7 x 4 x 1 cm consisting of 23 layers of an HMKON photographic emulsion from the P type was irradiated with slow positive pions of the CMMA (Ob"yedinenty, institut yadernykh issledovaniy - United Institute for Nuclear Research) synchrocyclotron. The chamber was mounted in a double magnetic screen in order to make sure that the scattered magnetic field of the synchrotron did not lead to a precession of the spin of the myon. In looking through the emulsions after developing those cases were selected where the whole myon track of the  $\pi \to \mu$ -decay

is situated in a single layer of the enulsion. In this the

Card 1/3

The Asympetry in the Angular Distribution of  $\mu^+ \Rightarrow e^+$  Decay 56-2-2/51 Electrons Observed in Photographic Enulsions

myon is supposed to come to a standstill after the passage through at least 50m of the surface of the non-developed layer of enulsion. The authors determined the angle a between the direction of emission of the myon in the  $\pi \to \mu-$ -decay and that of the electron of the  $\mu \rightarrow e$ -decay by determining the angle a between these directions on the emulsion level and the angle of sitribution  $\beta_1$ ,  $\beta_2$  resp. of the traces of the ayon, the electron towards the level of emulsion resp.. Furthermore an emulsion chamber of the same dimensions was irradiated with slow positive pions. The results of measurements are collected in a table. The angular distributions determined this way are shown by a diagram; they do not contradict the theoretical dependence 1 + a cose,  $a = (\lambda/3)(1 - \gamma)$ , where  $\gamma$  denotes the depolarization coefficient of myons. A relation for the determination of the optimum value of a is given. The magnetic field (H~1100 G) increases a little the asymmetry, i.e. it decreases the depolarization of the myons in the emulsion. But this effect is not regarded as strictly proved. The mean value of the parameter a calculated from the results of this work is  $a = -(0.108 \pm 0.0094)$ . The angular distribution for 13770 $\mu$ 

Card 2/3

The Asymmetry in the Angular Distribution of  $\mu^+ \rightarrow e^+$  Decay 56-2-2/51 Electrons Observed in Photographic Emulsions

decay processes proceeding from the results of various previous works and from those of the present investigation is also shown in a diagram. Within the error limits the angular distribution of the electrons of the relation  $1 + a \cos \theta$ , is sufficient, where  $a = -(0,111 \pm 0,015)$ . There are 2 figures, 2 tables, and 13 references, 1 of which is Slavic.

ASSOCIATION:

AS USSR (Akademiya nauk SSSR)

SUBMITTED:

August 14, 1957

AVAILABLE:

Library of Congress

1. Photographic emulsions-Irradiation 2. Klectrons-Distribution

Card 3/3

sov/56-36-4-65/70 Ali-Zade, S. A., Gurevich, I. I., Dobretsov, Yu. P., 24(3), 21(7) AUTHORS: Nikol'skiy, B. A., Surkova, L. V.

The Asymmetry of Electron Angular Distribution in At- e - Decay in a Magnetic Field of 27000 G (Asimmetriya uglovogo raspredeleniya TITLE: elektronov / et-raspada v magnitnom pole 27000 G)

Zhurnal eksperimental'noy i teoreticheskoy fiziki, 1959, Vol 36, PERIODICAL: Nr 4, pp 1327-1329 (USSR)

If angular distribution is described by the expression  $4\pi dN/do = 1-a \cos\theta (a = \lambda P/3 = a_0P; \lambda = 3a_0 = -\cos(V,A) ch...r$ ABSTRACT: acterizes the ratio of the vectorial and pseudovectorial share

of interaction in (1 - e-decay; P denotes muon polarization), it is found that the quantity a depends both on the measuring method and on the nature of the depolarized matter. It attains a maximum value of a = 1/3 at cos(A,V) = -1. For NIKFI-R emulsions a was determined as amounting to 0.092 ± 0.018, for Ilford G-5 it was 0.14. The maximum value attained by a for graphite is 0.303 ± 0.048. The depolarizing property of matter may be reduced by applying strong magnetic fields, the direction of which co-

incides with muon polarization. The increase of a brought about Card 1/3

SOV/56-36-4-65/70 The Asymmetry of Electron Angular Distribution in  $\mu^{+}$  e -Decay in a Magnetic Field of 27000 G

> by magnetic field can be described by  $a = a_0 \left[ 1 - \frac{0.5}{1 + (\mu H/\Delta E)^2} \right]$ ; a denotes the a-value if no depolarization takes place,  $\triangle E$  - the energy of fine-structure splitting of the u-mesic atom in the S-state. An experimental checking of this formula in fields of up to 14000 G showed that by it the dependence a(H) is qualitatively described. The authors determined a in the  $\pi$  - $\mu$ -edecay in photoemulsions at H = 27000 G. a was determined from the ratio a =  $2(N_{\text{backward}} - N_{\text{forward}})/(N_{\text{backward}} + N_{\text{forward}})$ . Results: For  $\theta = 0 - 30^{\circ}$   $a_1 = 0.315 \pm 0.026$  $a_2 = 0.295 \pm 0.027$ .  $\theta = 150 - 180^{\circ}$

> Mean value formation averaged over the directions in which muons fly off gives:  $a_3 = 0.305 \pm 0.019$ . If  $a_{real} = a_3/\cos\theta$ , one obtains  $a_{real} = a_3/0.940 = 0.324 \pm 0.020$ . Herefrom it follows that  $|\lambda|P = 0.972 \pm 0.06$ , i.e.  $|\lambda|$  with an accuracy of up to a

Card 2/3

The Asymmetry of Electron Angular Distribution in U+ e+Decay in a Magnetic Field of 27000 G

statistical error of  $^+$  6% attains its maximum value and P  $\approx$  1. This indicates a considerable degree of inaccuracy of the formula describing a(H). The authors finally thank B. S. Neganov and B. V. Sokolov for their help in irradiating the photoemulsions, D. M. Samoylovich for developing the emulsion, and further also V. M. Kutukova, A. M. Alpers, and G. V. Pleshivtseva for their assistance. There are 8 references, 2 of which are Soviet.

SUBMITTED:

February 1, 1959

Card 3/3

21 (7)

AUTHORS: Gurevich, I. I., Nikol'skiy, B. A.

SOV/56-37-1-58/64

Surkova L. V.

TITLE:

Three-Electron Decay of the u-Meson (Trekhelektronnyy raspad

µ-mezona)

PERIODICAL:

Zhurnal eksperimental noy i teoreticheskoy fiziki, 1959, Vol 37,

Nr 1, pp 318 - 319 (USSR)

ABSTRACT:

The authors of the present "Letter to the Editor", when investigating the asymmetry of the angular electron distribution of  $\pi \to \mu \to e\text{-decay}$ , observed that in one case three relativistic electrons depart from the stopping point of the  $\mu\text{-meson}$  (cf. figure). All three electrons have large inclination angles with respect to the emulsion plane, and therefore exact measurement of grain density was impossible; it was, however, near that for relativistic particles (energy-1 Mev). The recorded part of the electron path length:  $L_{e_1} = 455~\mu$ ,  $L_{e_2} = 562~\mu$ ,  $L_{e_3} = 455~\mu$ .

The muon range amounts to  $R_{ii} = 598 \mu$  in the case of an average

path length of the muon of the  $\pi \rightarrow \mu$ -decay in a R-NIKFI-emulsion of 602  $\mu$ . The angles between the electrons:  $\theta_{12} = 8.6^{\circ}$ ,  $\theta_{13} =$ 

Card 1/3

Three-Electron Decay of the u-Meson

807/56-37-1-58/64

as a muon decay in electron + f' with a transformation of the quantum into electron + positron at the place of decay, because in this case the direction of the departure of the particle pairs would have had to be opposite to the observed direction of flight of the decay electrons. The explanations  $\mu^+ \to e^+ + e^+ + e^- + p + p \to e^+ + p + p \to e^+ + p + p \to e^+ + p \to e^+$ 

Card 2/3

Three-Electron Decay of the µ-Meson

SOV/56-37-1-58/64

ing to 15 Mev. The authors finally thank I. S. Shapiro for dis-

cussions. There is 1 figure.

SUBMITTED:

April 17, 1959

Card 3/3

3/056/62/043/001/048/056 B102/B104

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Ivanov, Yu. M., Nikol'skiy, B. A., Smirnov, B. M.,

Surkova, L. V.  $\mu^*$ -meson depolarization in an electric field

TITLE

Zhurnal eksperimental noy i teoreticheskoy fiziki, v. 43,

no. 1(7), 1962, 337-339 era iodical:

TEXT: The authors studied the effect which a strong electric field  $(E \sim 10^5 \text{ V/cm})$  exerts on the depolarization of  $\mu^+$  mesons resulting from π-μ decays in photoemulsions. Depolarization of stopped muons is attributed mainly to production of muonium (µte); it has, however, also been observed (Swanson, Phys. Rev. 112, 580, 1958) that the "stopped" meson precessed in a transverse magnetic field and showed no further depolarization. Thus, muonium must be produced within a very short time immediately after the stoppage. It has not yet been verified by experiment whether the  $\mu^{+}$  meson in condensed matter decays as a free

Card 1/2

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particle or after having been cartured by a molecule. The authors found out that electrical fields of  $E\sim 10^5 v/cm$  were able to cause additional depolarization of the stopped muon; which is indicative of a captured muon. Such fields cannot depolarize muons in free or in muonium state. The experiments were made at the synchrocyclotron of the OIYaI with a palsed eg-Mev  $\pi^+$  beam (pulse duration 500  $\mu$ sec). The muons arising in a- $\mu$  decays were stopped in MNVVM-P (NIKFI-R) emulsions with increased gel time content, placed between electrodes. The pulsed field in the smulsion was  $1.2\cdot 10^5$  v/cm (400- $\mu$  emulsion layer) and  $2.4\cdot 10^5$  v/cm (200  $\mu$ ). The additional muon depolarization observed when the field was switched on proves that the stopped muon is captured by a gelatine molecule. There is 1 table.

SUB...ITTED: May 7, 1962

Card 2/2

Betainoniv E.C., Shin', V.G., Acciento, G.V., L'Khib, V.V., Surkova, L.V.

Monuros and simplina as algicides used against filamentous
algue. Tridy Gidrobiol. ob-va 14:52-65 '63. (MIRA 17:6)

1. Institut gidrobiologii AN UkrSSR, Kiyev.

L 3202-66 JXT(C2) ACC NR: AT5022299

SOURCE CODE: UR/3136/64/000/620/0001/0011

AUTHOR: Gurevich, I. I.; Makar'ina, L. A.; Nikol'skiy, B. A.; Sokolov, B. V.; Surkova, L. V.; Khakimov, S. Kh.; Shestakov, V. D.; Dobretsov, Yu. P.; Akhmanov, V. V.

ORG: [Gurevich, Makar'ina, Nikol'skiy, Sokolov, Surkova, Khakimov, Shestakov] IAE; [Dobretsov] MIFI; [Akhmanov] LYaP OIYaI

TITLE: Asymmetry of the angular distribution of electrons in the decay  $\pi^{+} + \mu^{+} + e^{+}$  in a magnetic field of 140,000 gauss

SOURCE: Moscow. Institut atomnoy energii. Doklady, IAE-620, 1964. Asimmetriya uglo-vogo raspredeleniya elektronov pi plus + mu plus + e plus raspada v magnitnom pole napryazhennost'yu 140 000 gauss, 1-11

TOPIC TAGS: mu meson, pi meson, positron, bubble chamber, radioactive decay

ABSTRACT: The universal V-A coupling theory applied to the determination of the angular distribution of electrons in the reaction  $\pi^+ + \mu^+ + e^+$  is given by

 $\frac{dN}{d\theta} \sim 1 - a \cos \theta_a$ 

in terms of the parameter a. In order to obtain a value of a which depends on the polarization state of the meson, an experiment was performed showing the effect countering the depolarization of the dense medium through which the meson is moving.

Card 1/2

L 10491-65 EST(1)/ENA(E) PEL GE/MAE

ACCESSION NR: AT4047629 S/0000/64/009/000/0339/0376

AUTHOR: Solov'yeva, S. F.: Sigachev, I. I.: Surkova, N. A.: Kogteva, Ye. V.

TITLE: Relay-and-microswitch-type contacts for small-signal switching

SOURCE. Vsesovisnove soveshchanive po elektricheskim kontaktam i

The second of t

TOPIC TAGS: small signal switching, small signal contact, microswitch contact, relay contact

ABSTRACT: An investigation of the effect of films on contact surfaces is presented to  $x_B > 0$  to the sweeth stated one Air-exposed contacts and (3) Sealed-in-plastic-container contacts. The deffect of long storaged was an inverse of the same o

Card 1/3

0

L 10441 65 ACCESSION NR: AT4047629

their alloys with Ni, Zr, Rh, were tested as contact materials. "Methods of less gating the electrophys cal characteristics of surface films formed on wta t majer sis included measuring the resistance, by voltmeter and ammeter, of a love a time seed the place spallings and a sphere made from Pt+ 25% Ir reserved the same of the same indentor was replaced which he is the sphere and the pressure of 0.5-200 g. was applied open-circuit voltage was 50 miv; current, 10-6 -10-1 amp. that there is a more than the property of the state of the pressure and the result of the result of the result of the state of the result of the reported of the was found with a relate in its operation in a modern that is a section of the section of the this take more site at electrical metaline ago, i.e. The best of them proved to be Aut with Process Associations of Aut 35 Ar deserve the attention of forther standard and of the Pt. March 19th 19th Rb can be recommended; however, we saily 4 times as high as that of Au+16% Pd . . . - in a exposed-contact designs. the same contacts in sealed designs was

Card 2/3

L 10491-65 ACCESSION NR: AT4047629

studied by placing both the plastic and contact material in a scaled envelope and subjecting it to thermal aging. Plastics AG-4S at 175C, K-211-3 and RST at 150C, and teffon at 250C were aged for 178 hrs. Dark spots were discovered on the metal after the aging test; the spots were due to the metal surface sorption of the metal after the aging test; the spots were due to the metal surface sorption of the volatile substances produced by the plastic materials. The spots on some of the metals and alloys offered a contact resistance up to tens of megohins. Silver the metals and alloys offered a contact resistance up to tens of megohins. Silver the metals and alloys offered a contact above plastics, showed a contact and the first the spots of the above plastics, showed a contact and the first the spots of the above plastics, showed a contact and the first the spots of the above plastics, and 3 tables.

ASSOCIATION: none

SUBMITTED: 13Jul64

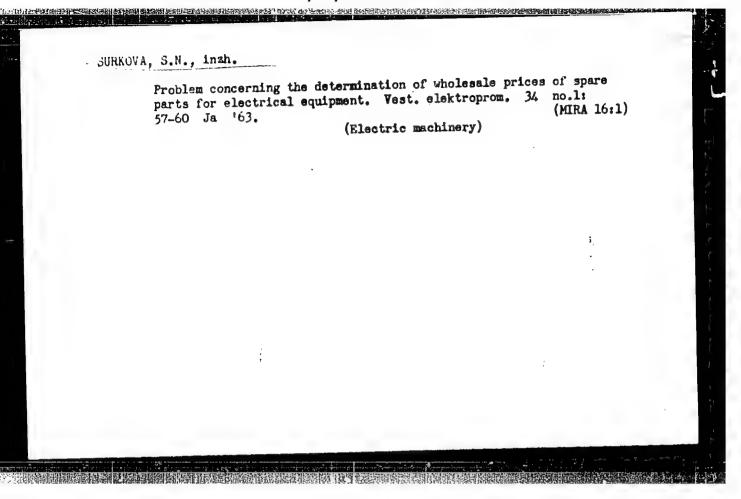
SUB CODE: EC, DP

NO REF SOV: 005

OTHER: 002

ENCL: 00

Card 3/3



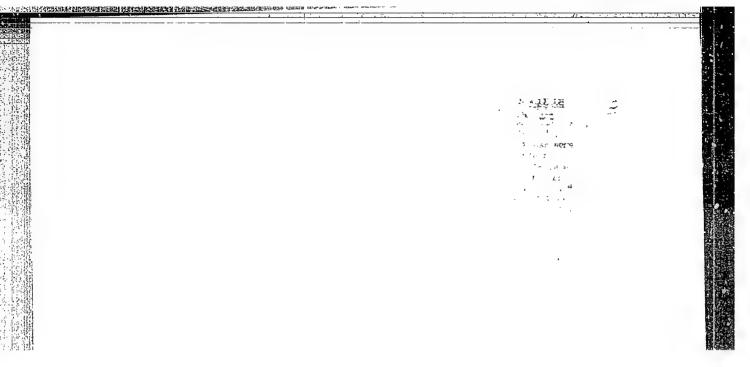
SURKOVA, S. H., inch.

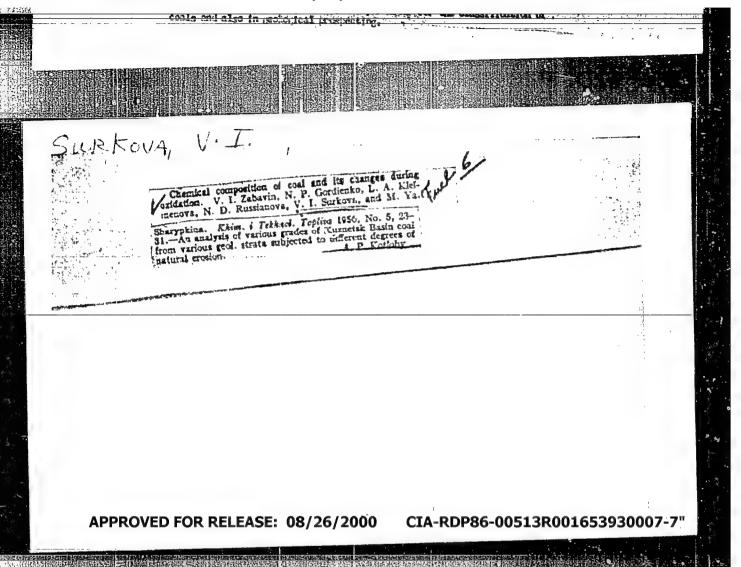
Economic expediency of electric motor repairs. Energ. i elektrotekh. prom. no.2:63-65 Ap-Je '64. (MRA 17:10)

SURROVA, S.N., inch.

Improved planning for the repair costs of electrical machines.
Energ. i elektrotekh. prom. no.2:64-66 Ap-Je \*65.

(MIRA 18:8)

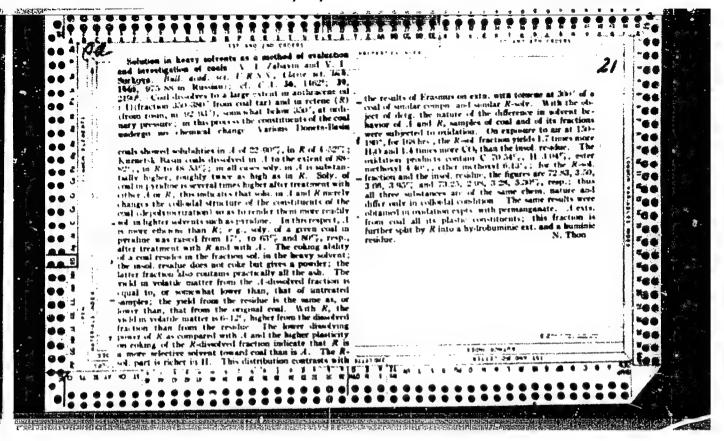


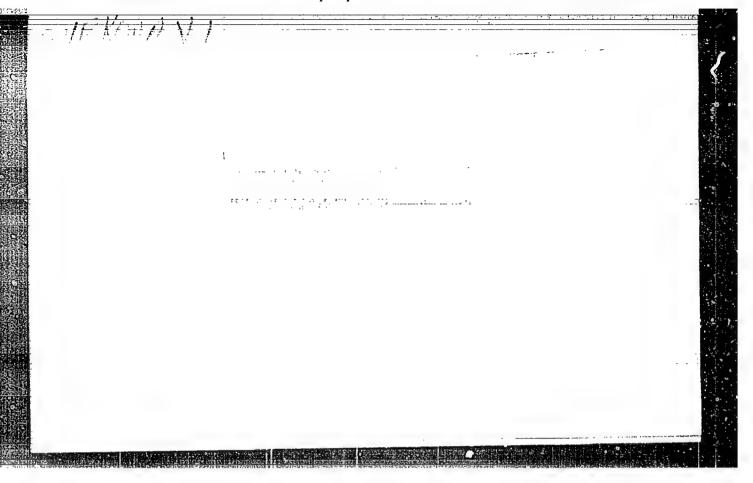


MASIEWNIKOV, E.N.; PIKOVSKAYA, O.G., starshly nauchnyy sotrudnik; SURKOVA, V.I., mladshiy nauchnyy sotrudnik; AGAFONOVA, L.L., mladshiy nauchnyy sotrudnik

Avivage preparations for polyvinyl alcohol fibers. Tekst. prom. 25 no.9:29-31 S \*65. (MIRA 18:10)

1. Rukovoditel' gruppy tekstil'noy persrabotki Vsesoyuznogo nauchno-issledovatel'skogo instituta iskusstvennogo volokna (for Maslennikov). 2. Laboratoriya otdelki i krasheniya Vsesoyuznogo nauchno-issledovatel'skogo instituta iskusstvennogo volokna (for Pikovskaya, Agafonova). 3. Gruppa tekstil'noy pererabotki Vsesoyuznogo nauchno-issledovatel'skogo instituta iskusstvennogo volokna (for Surkova).





# SURKOVA, V.L.

USSR/Chemical Technology - Chemical Products and Their Application. Treatment of solid mineral fuels

I-12

: Referat Zhur - Khiniya, No 4, 1957, 12838 Abs Jour

Zabavin V.I., Gordiyenko N.P., Kleymenova L.A., Russianova N.D., Surkova V.L., Sharypkina M Ya. Author

On Chemical Composition of Coal and Its Change on

Title Oxidation

Khimiya, i tekhnol. topliva, 1956, No 5, 23-31 Orig Fub

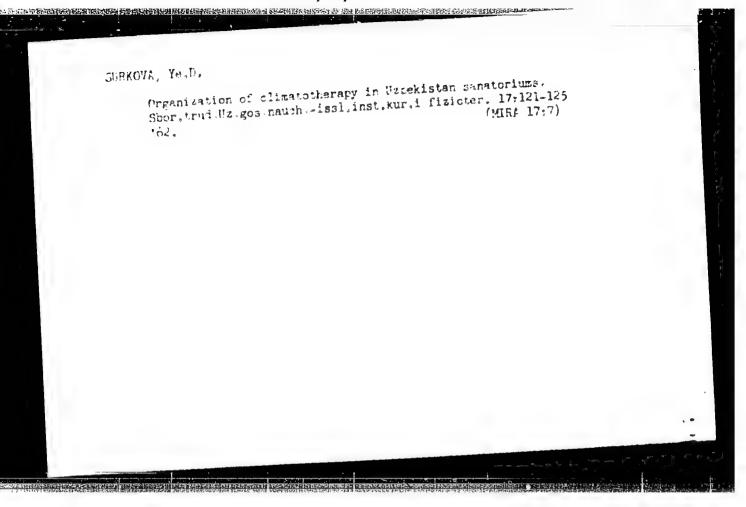
: Presented are the results of exhaustive "hot" extraction (in which the sample is heated by solvent vapor) of coal Abstract

of different grades from the new deposits of the Kuznetsk coal fields, unoxidized and of different degree of disintegration, with alcohol-benzene and with 5% solutions of KOH in alcohol-benzene removes from coal of grade D and U2 3-12% of extract, ~ 1% from coal of grade Zh, and > 0.5% from coal of grades K-TS. Yield of extract from

oxidized coal of grades G1 and Zh2 is higher than from

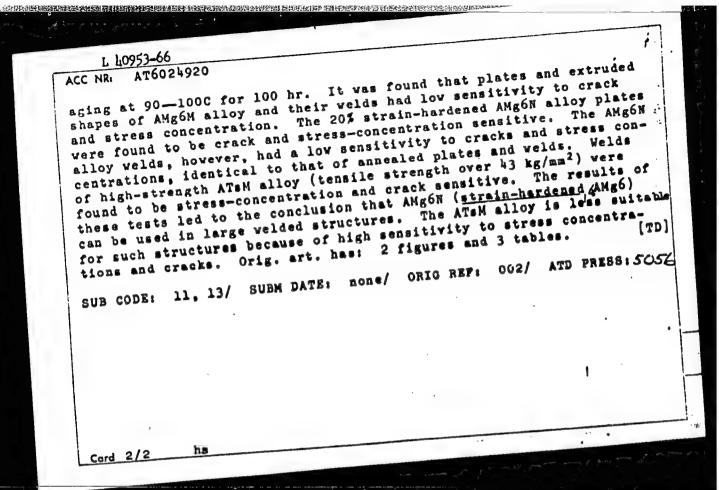
- 206 -Card 1/2

CIA-RDP86-00513R001653930007-7" APPROVED FOR RELEASE: 08/26/2000

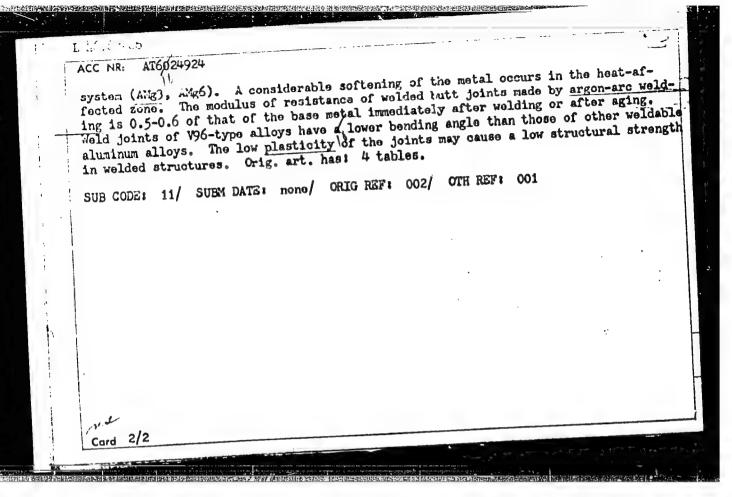


L 40753-00 EVT(m)/EVP(w)/EVP(k)/T/EVP(t)/ETI IJP(c) ACC NRI AT6024920 (A) SOURCE CODE: UR/2981/66/000/004/0112/0119 AUTHOR: Kishkina, S. I.; Zilova, T. K.; Kadobnova, N. V.; Drozdovskiy, e B. A.; Bubenshchikov, V. S.; Surkova, Yu. I. ORG: none TITLE: Stress-concentration and crack sensitivity of ATSM, ATSMU and AMG6, alloys and their welds SOURCE: Alyuminiyevyye splavy, no. 4, 1966. Zharoprochnyye i vysokoprochayye splayy (Heat-resistant and high-strength alloys), 112-119) TOPIC TAGS: "aluminum alloy, high strength alloy, stress concentration, notch sensitivity, metal property, / ATSM aluminum alloy, ATSMU aluminum alloy, AMg6M aluminum alloy, AMg6N aluminum alloy ABSTRACT: Hot-rolled ATSM, ATSMU, AMg6M and AMg6N alloy plates 10 mm thick, ATSM and ATSMU alloy forgings, ATSMU and AMg6M alloy extruded shapes, and welds of these alloys have been tested for stress-concentration and crack sensitivity. The sensitivity to stress concentratition was evaluated on the basis of tensile tests with notched specimens stressed under an angle of 4-8° to the axis. Crack sensitivity was tested with Mesnoger specimens having artificial cracks 1.5 mm deep. In all cases, specimens of ATSM and ATSMU alloys were tested after Card 1/2

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L 46968-66 EMP(k)/EMT(m)/T/EMP(w)/EMP(v)/EMP(t)/ETI IJP(c) JH/JD/HS SOURCE CODE: UR/2981/66/000/004/0152/0158	P <sub>1</sub>
ACC NR: AT6024924 CF1147	1
Nei Vlasova, T. A.; Skachkov, III. I.; Dzyubenko, M. V.	
AUTHOR: Fridlyander, I. N.; Vlasova, T. A.; Skachkov, Yu. N.; Shillyander, M. V. Surkova, Yu. I.; Gorokhova, T. A.; Ped, A. A.; Gur, yev, I. I.; Dzyubenko, M. V.	
Surkova, Yu. 1.; Goroknova, 246	
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SOURCE: Alyuminiyevyye splavy, no. 4, 1, 152-158 (Heat resistant and high-strength alloys), 152-158	
(Heat resistant and high-strength alloys), 132 of the control of t	
TOPIC TAGS: Bluminum same	
zinc alloy  ABSTRACT: The object of the work was to study the weldability in the fusion welding  ABSTRACT: The object of the work was to study the weldability of this alloy can be im-	
ABSTRACT: The object of the work was to study the weldability in the Iusian was abe imported by and also to determine whether the weldability of this alloy can be imported by changing the chemical composition of the base metal and filler wire. Sheet proved by changing the chemical composition 8.44% In, 2.72% Mg, 2.2% Cu, proved by changing the chemical composition 8.44% In order to	5
of Vich alloy, and allow a series on of the base were a god ve 2 25 Cile	
of V96 alloy 2.5 mm thick of the chemical composition in the experiments. In order to 0.06% Mn, 0.13% Zr, 0.29% Fe, and 0.13% Si were used in the experiments. In order to 0.06% Mn, 0.13% Zr, 0.29% Fe, and 0.13% Si were used in the experiments should be carried ou decrease the tendency toward crystallization cracks) the welding should be carried ou decrease the tendency toward crystallization cracks) the welding should be carried ou decrease the tendency toward crystallization cracks. The content of the main alloying elements in the Al-Mg alloy fillers (of type AMg6). The content of the main alloying elements in the Al-Mg alloy fillers (of type AMg6).	n
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是一个人,我们就是这个人的,他们们也不是一个人的,你们不会的一个人的,你就是我们的一个人的人的人的人的人的人的人的人的人的人,我们也是我们的人,我们就是我们的人

CIBIAAS, F., kand. med. nauk; DAKTAMAVICIENE, E., kand. med. nauk;

JARZEMSKAS, J., kand. med. nauk [deceased]; JOCEVICIENE, A.,

kand. med.nauk; KRIKSTOFAITIS, M., kand. med. nauk; NENISKIS, J.,

kand. med. nauk; STEPONATTIENE, L., kand. med. nauk; SURKUS, J.,

kand. med. nauk; SIMANAS, S., kand. biolog. nauk; CEPULIS, St.,

prof.; KUPCINSKAS, J., prof.; LASAS, Vl., prof.; SIDERAVICIUS, Br.,

prof.; KANOPKA, E., dots.; KVIKINS, V., dots.; LABANAUSKAS, K.,

prof.; KANOPKA, E., dots.; BABUBINS, P., doktor; CAPKEVICIUS, V.,

dots.; POLUKO.DAS, H., dots.; BABUBINS, P., doktor; STUCKA.R.,

doktor; NAKARIUNAS, P., doktor; PAULIUKONIENE, J., red.; ANAITIS, J.,

tekhn. red.

[Health and diseases] Antrasis pataisytas leidimas. Vilnius, Valstybine politimes ir mokslines literaturos leidykla, 1961. 356 p. (MIRA 15:3)

(HYGIENE) (PATHOLOGY)

LAPYTE, A.; SURKUS, J., med. m. kand.

On the differential diagnosis of schizophrenia and mental disorders in theumatism. Sveik. apsaug. 8 no.4:25-29 Ap<sup>1</sup>63.

1. Resp. Kauno psichoneurologine ligonine ir Kauno Valst. medicinos institutas.

SURKUS, J., doc. med. m. kand.

The psychoneurological hospital of the Republic of NauyaVilna (on the 60th anniversary of its founding). Syeik. apsaug.

8 no.5:32-37 163.

1. Kauno Valst. medicinos institutas.
(HOSPITALS) (HOSPITALS, PSYCHIATRIC)
(NEUROLOGY)

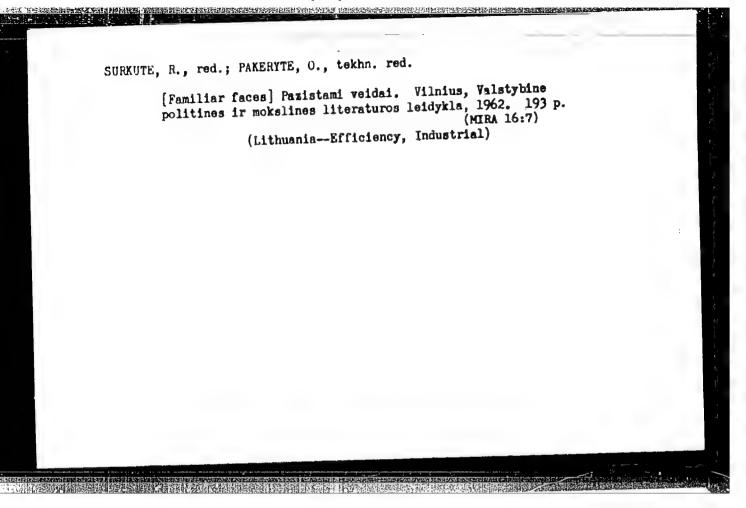


Fig. 7. The secondary subtree and leukoplastic injury to the edges of wounds, New Jobs, post sent in-ta, Ai. II, 19., p. 183-49.

3.: U-338, 19 Aurust 53, (Letopis 'neki Statey', No. 22, 1989).

OFRIA, Henry:: FAM: ILDERI, Maria

Apreces of the role of adrenals in fengetherapy of nomen with postinflamatory adnexan changes and sterility. Sinck. sel. 35 no. 1.22/-

237 Mr-Ap 164.

1. E 11 Kliniki Poloznictwa i Chorca kobiscych Akademii Madyranej w Głensku (Klerownik: Prof. dr. nei. m. Grennicki)

作的大型的工程,但是由于1850年的,1950年,1950年,1950年,1950年,1950年,1950年,1950年,1950年,1950年,1950年,1950年,1950年,1950年,1950年,1950年,1

PANKISS, T, Marie; SUMMA, Henryk

Offect of believelogical therapy on the exerction of 17-ketrateroids, 17- hydroxycerticosteroids and total steroids in the urine of sexually mature women before and after surgical removal of cvaries. Ginek. Pol. 35 nc.4:561-567 Jl- Ag \*64.

1. Z II Klinika Foloznictwa i Chorob Kobiecych Akademii Medycznej w Gdansku i z Osrodka Naukowo Badawczego wFolozynie-Zdroju (Kierownik: prof. dr. med. W. Gromadzki).

Exerction of metabolites of the advanal cortex curing treatment of adnexitis with glycocorticoids. Ginek. Fol. 36 no.9:999-1005 S '65.

1. 2 II Kliniki Poloznictwa i Chorob Koblecych AM w Gdansku (Kierownik: prof. dr. med. W. Gromadzki).

P/518/62/011/001/004/008 D207/D308

AUTHOR:

Surma, Marian

TITLE:

Magnetic birefringence of liquid solutions. I. Magnetic birefringence of polar liquids and their solutions

in carbon tetrachloride

SOURCE:

Poznańskie Towarzystwo Przyjaciół Nauk. Komisja

Matematyczno-Przyrodnicza. Prace. v. 11, no. 1, 1962,

Fizyka dielektryków. no. 1, 113 - 128

TEXT: This paper was presented on October 19, 1961 at a meeting of the Komisja Matematyczno-Przyrodnicza PTPN (Mathematical and Scientific Committee, PTPN). The Cotton-Mouton magnetic birefringence constants, C, of nitrobenzene, m-nitrotoluene, o-nitrotoluene, o-nitroanisol, 1,2,-dibromoethane and their solutions in carbon tetrachloride were measured at room temperature using light of  $\lambda = 547$  mm. All these liquids were purified by distillation and (except CCl<sub>4</sub>) by filtering through Al<sub>2</sub>O<sub>3</sub>. The constants were deduced from the phase difference between the ordinary and extraordinary rays Card 1/2